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On the stability of the massive scalar field in Kerr space-time

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The current early stage in the investigation of the stability of the Kerr metric is characterized by the study of appropriate model problems. Particularly interesting is the problem of the stability of the solutions of the Klein-Gordon equation, describing the propagation of a scalar field in the background of a rotating (Kerr-) black hole. Results suggest that the stability of the field depends crucially on its mass $\mu$. Among others, the paper provides an improved bound for $\mu$ above which the solutions of the reduced, by separation in the azimuth angle in Boyer-Lindquist coordinates, Klein-Gordon equation are stable. Finally, it gives new formulations of the reduced equation, in particular, in form of a time-dependent wave equation that is governed by a family of unitarily equivalent positive self-adjoint operators. The latter formulation might turn out useful for further investigation. On the other hand, it is proved that from the abstract properties of this family alone it cannot be concluded that the corresponding solutions are stable.

I. INTRODUCTION

Kerr space-time is the only possible vacuum exterior solution of Einstein’s field equations describing a stationary, rotating, uncharged black hole with non-degenerate event horizon and is expected to be the unique, stationary, asymptotically flat, vacuum space-time containing a non-degenerate Killing horizon. Also, it is expected to be the asymptotic limit of the evolution of asymptotically flat vacuum data in general relativity.

An important step towards establishing the validity of these expectations is the proof of the stability of Kerr space-time. In comparison to Schwarzschild space-time, this problem is complicated by a lower dimensional symmetry group and the absence of a Killing field that is everywhere timelike outside the horizon. For instance, the latter is reflected in the fact that energy densities corresponding to the Klein-Gordon field in a Kerr gravitational field have no definite sign. This absence complicates the application of methods from operator theory and of so-called “energy methods” that are both employed in estimating the decay of solutions of hyperbolic partial differential equations.

On the other hand, two facts are worth noting. For this, note that in the following any reference to coordinates implicitly assumes use of Boyer-Lindquist coordinates.

First, in addition to its Killing vector fields that generate one-parameter groups of symmetries (isometries), Kerr space-time admits a Killing tensor that is unrelated to its symmetries. Initiated by his groundbreaking work on the complete separability of the Hamilton-Jacobi equation in a Kerr background, Carter discovered that an operator that is induced by this Killing tensor commutes with the wave operator. On the other hand, Carter’s operator contains a second order time derivative. An analogous operator has been found for the operator governing linearized gravitational perturbations of the Kerr geometry. A recent study finds another such “symmetry operator” which only contains a first order time derivative and commutes with a rescaled wave operator. Differently to Carter’s operator, this operator is analogous to symmetry operators induced by one-parameter group of isometries of the metric, in that it induces a mapping in the data space that is compatible

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with time evolution and, therefore, describes a true symmetry of the solutions. It is likely that an analogous operator can be found for a rescaling of the linearized operator governing gravitational perturbations of the Kerr geometry. In case of existence, it should facilitate the generalization to a Kerr background of the Regge-Wheeler-Zerilli-Moncrief (RWZM) decomposition of fields on a Schwarzschild background, which in turn should greatly simplify the analysis of the stability of Kerr space-time.

Second, there is a Killing field that is timelike in an open neighborhood of the event horizon given by

$$\xi := \partial_t + \frac{a}{2Mr_+} \partial_\varphi,$$

where \(\partial_t, \partial_\varphi\) are coordinate vector fields of Boyer-Lindquist coordinates corresponding to the coordinate time \(t\) and the azimuthal angular coordinate \(\varphi\). \(M > 0\) is the mass of the black hole and \(a \in [0, M]\) its rotational parameter. Moreover, if

$$\frac{a}{M} \leq \frac{\sqrt{3}}{3},$$

\(\xi\) is timelike in the ergoregion, see Lemma 2.2. On the other hand, \(\partial_t\) itself is spacelike in the ergoregion, null on the stationary limit surface, and timelike outside. For these reasons, at least for \(a\) satisfying (1.2), it might be possible to “join” energy inequalities belonging to the Killing fields by \(\xi\) and \(\partial_t\).

The discussion of the stability of the Kerr black hole is in its early stages. The first intermediate goal is the proof or disproof of its stability under “small” perturbations. In the case of the Schwarzschild metric, by using the RWZM decomposition of fields in a Schwarzschild background, the question of the stability can be completely reduced to the question of the stability of the solutions of the wave equation on Schwarzschild space-time. For Kerr space-time, a similar reduction is not known. If such reduction exists, there is no guarantee that the relevant equation is the scalar wave equation. It is quite possible that such equation contains an additional (even positive) potential term that, similar to the potential term introduced by a mass of the field, could result in instability of the solutions. Second, an instability of a massive scalar field in a Kerr background could indicate instability of the metric against perturbations by matter which generically has mass. If this were the case, even a proof of the stability of Kerr space-time could turn out as a purely mathematical exercise with little relevance for general relativity. Currently, the main focus is the study of the stability of the solutions of the Klein-Gordon field on a Kerr background with the hope that the results lead to insight into the problem of linearized stability. Although the results of this paper also apply to the case that \(\mu = 0\), its main focus is the case of Klein-Gordon fields of mass \(\mu > 0\).

Quite differently from the case of a Schwarzschild background, the results for these test cases suggest an asymmetry between the cases \(\mu = 0\) and \(\mu \neq 0\). In the case of the wave equation, i.e., \(\mu = 0\), results point to the stability of the solutions, whereas for \(\mu \neq 0\), there are a number of results pointing in the direction of instability of the solutions under certain conditions.

In particular, unstable modes were found by the numerical investigations by Furushashi and Nambu for \(\mu M \sim 1\) and \((a/M) = 0.98\), by Strafuss and Khanna for \(\mu M \sim 1\) and \((a/M) = 0.9999\), and by Cardoso and Yoshida for \(\mu M \leq 1\) and \(0.98 \leq (a/M) < 1\). The analytical study by Hod and Hod finds unstable modes for \(\mu M \sim 1\) with a growth rate which is four orders of magnitude larger than previous estimates. On the other hand, Ref. 3 proves that the restrictions of the solutions of the separated, in the azimuthal coordinate, Klein-Gordon field (RKG) are stable for

$$\mu \geq \frac{|m|a}{2Mr_+} \left(\frac{2M}{r_+} + \frac{a^2}{r_+^2}\right).$$

Here, \(m \in \mathbb{Z}\) is the “azimuthal separation parameter” and \(r_+ := M + \sqrt{M^2 - a^2}\). So far, this has been the only mathematically rigorous result on the stability of the solutions of the RKG for \(\mu > 0\). This result contradicts the result of Zouros and Eardley, but is consistent with the other results above. In addition, there is the numerical result by Konoplya and Zhidenko, which confirms the result of Beyer, but also finds no unstable modes of the RKG for \(\mu M \ll 1\) and \(\mu M \sim 1\).
Among others, this paper improves the estimate (1.3). It is proved that the solutions of the RKG are stable for $\mu$ satisfying

$$\mu \geq \frac{|m|a}{2Mr_+} \sqrt{1 + \frac{2M}{r_+}}.$$ 

Further, it gives new formulations for RKG, in particular, in form of a time-dependent wave equation that is governed by a family of unitarily equivalent positive self-adjoint operators. The latter might turn out useful in future investigations. On the other hand, it is proved that from the abstract properties of this family alone it cannot be concluded that the corresponding solutions are stable.

The remainder of the paper is organized as follows. Section II gives the geometrical setting discussed in the solutions of the RKG and a proof of the above-mentioned property of the Killing field $\xi$. Section III gives basic properties of operators read off from the equation, including some new results. These properties provide the basis for a formulation of the initial-value problem for the equation in Sec. IV, which is less dependent on methods from semigroups of operators than that of Ref. 3. Section IV also contains the improved result on the stability of the solutions of RKG, a formulation of the RKG in terms of a time-dependent wave equation and the above-mentioned counterexample. Finally, the paper concludes with a discussion of the results and two Appendixes that contain proof of results that were omitted in the main text to improve the readability of the paper.

II. THE GEOMETRICAL SETTING

In Boyer-Lindquist coordinates, $\Omega \rightarrow \mathbb{R}^4$, the Kerr metric $g$ is given by

$$g = g_{tt} dt \otimes dt + g_{\varphi \varphi} (dt \otimes d\varphi + d\varphi \otimes dt) + g_{rr} dr \otimes dr + g_{\theta \theta} d\theta \otimes d\theta + g_{\varphi \theta} d\varphi \otimes d\theta,$$

where

$$g_{tt} := 1 - \frac{2Mr}{\Sigma}, \quad g_{\varphi \varphi} := \frac{2Mar \sin^2 \theta}{\Sigma}, \quad g_{rr} := -\frac{\Delta}{\Sigma}, \quad g_{\theta \theta} := -\frac{\Delta}{\Sigma} \sin^2 \theta,$$

$M$ is the mass of the black hole, $a \in [0, M]$ is the rotational parameter and

$$\Delta := r^2 - 2Mr + a^2, \quad \Sigma := r^2 + a^2 \cos^2 \theta,$$

$$\Delta := \frac{(r^2 + a^2)^2 \Sigma + 2Ma^2r \sin^2 \theta}{\Sigma} = \frac{(r^2 + a^2)^2}{\Delta} - a^2 \sin^2 \theta = \Sigma + 2Mr + \frac{4M^2r^2}{\Delta},$$

$$r_+ := M + \sqrt{M^2 - a^2}, \quad r_- := M - \sqrt{M^2 - a^2},$$

$$\Omega := \mathbb{R} \times (r_+, \infty) \times (0, \pi) \times (-\pi, \pi).$$

In these coordinates, the reduced Klein-Gordon equation corresponding to $m \in \mathbb{Z}$, governing solutions $\psi : \Omega \rightarrow \mathbb{C}$ of the form

$$\psi(t, r, \theta, \varphi) = \exp(im\varphi)u(t, r, \theta),$$

where $u : \Omega_t \rightarrow \mathbb{C}$,

$$\Omega_t := (r_+, \infty) \times (0, \pi),$$

for all $t \in \mathbb{R}, \varphi \in (-\pi, \pi), (r, \theta) \in \Omega$, is given by

$$\frac{\partial^2 u}{\partial t^2} + ib \frac{\partial u}{\partial t} + D_{\varphi \theta}^2 u = 0,$$

where

$$b := \frac{4mMar}{\Delta \Sigma_0} = \frac{4mMar}{(r^2 + a^2)^2 - a^2 \Delta \sin^2 \theta} = \frac{4mMar}{(r^2 + a^2)^2 \Sigma + 2Ma^2r \sin^2 \theta},$$

$$D_{\varphi \theta}^2 f := \frac{1}{\Sigma_0} \left( -\frac{\partial}{\partial r} \frac{\partial}{\partial r} - \frac{m^2a^2}{\Delta} - \frac{1}{\sin \theta} \frac{\partial}{\partial \theta} \sin \theta \frac{\partial}{\partial \theta} + \frac{m^2}{\sin^2 \theta} + \mu^2 \Sigma \right) f.$$
for every \( f \in C^2(\Omega_s, \mathbb{C}) \) and \( \mu \geq 0 \) is the mass of the field. In particular, note that \( b \) defines a real-valued bounded function on \( \Omega_s \), which is positive for \( m \geq 0 \) and negative for \( m \leq 0 \). For this reason, it induces a bounded self-adjoint (maximal multiplication) operator \( B \) on the weighted \( L^2 \)-space \( X \), see below, which is positive for \( m \geq 0 \) and negative for \( m \leq 0 \). Further, \( D_{r_0}^2 \) is singular since the continuous extensions of the coefficients of its highest (second) order radial derivative vanish on the horizon \( \{ r_+ \} \times [0, \pi] \).

In particular, the following proves that the Killing field

\[
\xi := \partial_t + \frac{a}{2Mr_+} \partial_\varphi
\]

is timelike in an open neighborhood of the event horizon and timelike in the ergoregion if

\[
\frac{a}{M} \leq \frac{\sqrt{3}}{3}.
\]

Proofs are given in Appendix A.

**Lemma 2.1:** Let \( M > 0, a > 0 \). For every \( s \in \mathbb{R} \), the function

\[
g(\partial_t + s \partial_\varphi, \partial_t + s \partial_\varphi)
\]

has a continuous extension to \( \Omega_s \). This extension is positive on \( \partial \Omega_s \) if and only if

\[
s = \frac{a}{2Mr_+}.
\]

Further,

\[
\xi := \partial_t + \frac{a}{2Mr_+} \partial_\varphi
\]

is timelike precisely on

\[
\Omega_{e2} := \left[2Mr_+ - a^2 \sin^2 \theta - a \Delta^{1/2} \sin \theta \left(1 + \frac{2M}{r - r_-}\right)\right]^{-1}((0, \infty)).
\]

**Proof:** See Appendix A.

**Lemma 2.2:** Let \( M > 0, a > 0 \), and \( \Omega_{e1}, \) defined by

\[
\Omega_{e1} := (a^2 \sin^2 \theta - \Delta)^{-1}((0, \infty)),
\]

denote the ergoregion. If

\[
\frac{a}{M} \leq \frac{\sqrt{3}}{3},
\]

then

\[
\Omega_{e1} \subset \Omega_{e2}.
\]

**Proof:** See Appendix A.

### III. BASIC PROPERTIES OF OPERATORS IN THE EQUATION

In a first step, we represent (2.1) as a differential equation for an unknown function \( u \) with values in a Hilbert space. For this reason, we represent formal operators present in (2.1) as operators with well-defined domains in an appropriate Hilbert space and, subsequently, study basic properties of the resulting operators. Theorems 3.5 and 3.6 provide new results.

**Definition 3.1:** In the following, \( X \) denotes the weighted \( L^2 \)-space \( X \) defined by

\[
X := L^2_2(\Omega_\varphi, \Sigma_b \sin \vartheta).
\]

Further, \( B \) is the bounded linear self-adjoint operator on \( X \) given by

\[
Bf := bf
\]
for every \( f \in X \). Note that \( B \) is positive for \( m \geq 0 \) and negative for \( m \leq 0 \).

**Remark 3.2:** We note that, as consequence of the fact that \( B \in L(X, X) \) is self-adjoint, the operator

\[
\exp((it/2)B),
\]

where \( \exp \) denotes the exponential function on \( L(X, X) \), see, e.g., Sec. 3.3 in Ref. 5, is unitary for every \( t \in \mathbb{R} \) and coincides with the maximal multiplication operator by the function \( \exp((it/2)b) \).

**Definition 3.3 (Definition of \( A_0 \))**:

(i) We define \( D(A_0) \) to consist of all \( f \in C^2(\Omega_+, \mathbb{C}) \cap X \) satisfying the following conditions (a), (b), and (c):

(a) \( D_{r, \theta}^2 f \in X \),

(b) there is \( R > 0 \) such that \( f(r, \theta) = 0 \) for all \( r > R \) and \( \theta \in I_0 := (0, \pi) \),

(c) \[ \lim_{r \to r_0} \frac{\partial f}{\partial \theta}(r, \theta) = 0 \]

for all \( \theta \in I_0 \).

(ii) For every \( f \in D(A_0) \), we define

\[ A_0 f := D_{r, \theta}^2 f. \]

**Lemma 3.4:** \( A_0 \) is a densely defined, linear, symmetric, and essentially self-adjoint operator in \( X \). In addition, the closure \( \bar{A}_0 \) of \( A_0 \) is semibounded with lower bound

\[ \alpha := -\frac{m^2 a^2}{4 M^2 r_0^2}. \]

**Proof:** See Lemma 2 and Theorem 4 in Ref. 6.

**Theorem 3.5:** The span, \( D \), of all products

\[ f \otimes (P^m_l \circ \cos), \]

where \( f \in C^2_0((r_+, \infty), \mathbb{C}) \) and \( P^m_l : (-1, 1) \to \mathbb{R} \) is the generalized Legendre polynomial corresponding to \( m \in \mathbb{Z} \) and \( l \in \{ |m|, |m| + 1, \ldots \} \), is a core for \( \bar{A}_0 \).

**Proof:** For this, we use the notation of Ref. 6. According to the proof of Theorem 4 of Ref. 6, the underlying sets of \( X \) and \( \overline{X} := L^2(\Omega_+, (r^4/\Delta \sin \theta)) \) are equal; and the norms induced on the common set are equivalent, the maximal multiplication operator \( T_{r^4/\Delta \Sigma_0} \) by the function \( r^4/\Delta \Sigma_0 \) is a bijective bounded linear operator on \( X \) that has a bounded linear inverse; the operator \( H \), related to \( A_0 \) by

\[ A_0 = T_{r^4/\Delta \Sigma_0} H \quad (3.3) \]

is a densely defined, linear, symmetric, semibounded, and essentially self-adjoint operator in \( \overline{X} \), and \( D \) is contained in the (coinciding) domains of \( A_0 \) and \( H \). Further, it has been shown that \( (H - \lambda)D \) is dense in \( \overline{X} \) for \( \lambda < \beta \), where \( \beta := -m^2 a^2/r_0^4 \) is a lower bound for \( H \). From this follows that \( D \) is a core for the closure \( \overline{H} \) of \( H \). For the proof, let \( f \in D(\overline{H}) \). Since \( (H - \lambda)D \) is dense in \( \overline{X} \), there is a sequence \( f_1, f_2, \ldots \) in \( D \) such that

\[ \lim_{v \to \infty} (H - \lambda)f_v = (\overline{H} - \lambda)f. \]

Since \( \overline{H} - \lambda \) is bijective with a bounded inverse, the latter implies that \( f_1, f_2, \ldots \) is convergent to \( f \) and also that

\[ \lim_{v \to \infty} Hf_v = \overline{H}f. \]
Hence, we conclude that $\tilde{H}$ coincides with the closure of $H_{[D]}$. Since $T_{r^j(\Delta \Sigma_0)}, T_{r^{-j}(\Delta \Sigma_0)} \in L(X, X)$, from the latter also follows that $A_0$ coincides with the closure of $A_0|_{[D]}$. \hfill \qed

**Theorem 3.6:** The operator $A_0$ coincides with the Friedrichs extension of the restriction of $A_0$ to $C_0^\infty(\Omega_x, C)$. \hfill \qed

**Proof:** As a consequence of Theorem 3 in Ref. 3, it follows that $D$ is contained in the domain of the Friedrichs extension $\bar{A}_F$ of the restriction of $A_0$ to $C_0^\infty(\Omega_x, C)$ and that $A_F f = A_0 f$ for every $f \in D$. In this connection, note that the addition of a multiple of the identity operator “does not affect” the Friedrichs extension of an operator.\(^{37}\) Since $D$ is a core for $A_0$, from this follows that $A_F \supset A_0$ and hence, since $A_F$ is in particular symmetric and $\bar{A}_0$ is self-adjoint, that $A_F = \bar{A}_0$. \hfill \qed

**Lemma 3.7:** \hfill \qed

\[ \mathcal{A} := \bar{A}_0 + \alpha (1/4) B^2 \]

is a densely defined, linear, and positive self-adjoint operator in $X$.

**Proof:** That $\mathcal{A}$ is a densely defined, linear, and self-adjoint operator in $X$ is a consequence of Theorem 3.1 and the Rellich-Kato theorem. For the latter, see, e.g., Theorem X.12 in Ref. 32, Vol. II. The positivity of $\mathcal{A}$ is a simple consequence of the fact that

\[ \frac{1}{\Sigma_b} \left( - \frac{m^2 a^2}{\Delta} + \frac{m^2}{\sin^2 \theta} \right) + \frac{4 b^2}{\Delta} = m^2 \left[ \frac{\Delta - a^2 \sin^2 \theta}{\Delta \Sigma_b \sin^2 \theta} + \frac{4 M^2 a^2 r^2}{(\Delta \Sigma_b)^2} \right] \]

\[ = \frac{m^2}{(\Delta \Sigma_b)^2 \sin^2 \theta} \left[ (\Delta - a^2 \sin^2 \theta) \Delta \Sigma_b + 4 M^2 a^2 r^2 \sin^2 \theta \right] \]

\[ = \frac{m^2}{(\Delta \Sigma_b)^2 \sin^2 \theta} \left[ (\Delta - a^2 \sin^2 \theta) (\Delta \Sigma + 2 M r) + 4 M^2 a^2 r^2 \sin^2 \theta \right] \]

\[ = \frac{m^2}{(\Delta \Sigma_b)^2 \sin^2 \theta} \left[ (\Sigma - 2 M r) (\Sigma + 2 M r) + 4 M^2 r^2 \right] \]

\[ = \frac{m^2}{(\Sigma - 2 M r) (\Sigma + 2 M r) + 4 M^2 r^2} \geq 0. \]

\hfill \qed

**IV. FORMULATION OF AN INITIAL VALUE PROBLEM**

In the following, we give an initial value formulation for equations of the type of (2.1) whose possibility is indicated by Theorem 4.11 in Ref. 4, see also Theorem 5.4.11 in Ref. 5. Here, we give the details of such formulation, including abstract energy estimates that provide an independent basis for the estimate (1.3) and also for its improvement (4.5) below. Specialization of the abstract formulation to $X$ given by (3.1), $A := \tilde{A}_0 - \tilde{C}, B$ given by (3.2) and $\tilde{C} := -(\alpha + \epsilon)$ for some $\epsilon > 0$, provides an initial-value formulation for (2.1) on every open interval $I$ of $\mathbb{R}$ along with quantities that are conserved under time evolution. Note that in this case $A + C = \bar{A}_0$.

**Assumption 4.1:** In the following, let $(X, \langle \cdot | \cdot \rangle)$ be a non-trivial complex Hilbert space and $A$ be a densely defined, linear, and strictly positive self-adjoint operator in $X$.

**Definition 4.2:** We denote by $W^1_A$ the complex Hilbert space given by $D(A^{1/2})$ equipped with the scalar product $\langle | \rangle_1$, defined by

\[ \langle \xi | \eta \rangle_1 := \langle A^{1/2} \xi | A^{1/2} \eta \rangle + \langle \xi | \eta \rangle \]

for every $\xi, \eta \in D(A^{1/2})$, and induced norm $\| |$.

**Remark 4.3:** Note that, as a consequence of

\[ \| \xi \|_1 = (\| A^{1/2} \xi \|^2 + \| \xi \|^2)^{1/2} \geq \| \xi \| \]

\[ \| \xi \|_1 = (\| A^{1/2} \xi \|^2 + \| \xi \|^2)^{1/2} = \| \xi \| \]

\[ \| \xi \|_1 = (\| A^{1/2} \xi \|^2 + \| \xi \|^2)^{1/2} \geq \| \xi \| \]

\[ \| \xi \|_1 = (\| A^{1/2} \xi \|^2 + \| \xi \|^2)^{1/2} \geq \| \xi \| \]

\[ \| \xi \|_1 = (\| A^{1/2} \xi \|^2 + \| \xi \|^2)^{1/2} \geq \| \xi \| \]
for every $\xi \in D(A^{1/2})$, the imbedding $W^1_A \hookrightarrow X$ is continuous.

Assumption 4.4: Let $B : D(A^{1/2}) \to X$ be a symmetric linear operator in $X$ for which there are $a \in [0, 1)$ and $b \in [0, \infty)$ such that

$$\|B\xi\|^2 \leq a^2 \|A^{1/2}\xi\|^2 + b^2 \|\xi\|^2$$

for every $\xi \in D(A^{1/2})$. Note that this implies that $B \in L(W^1_A, X)$. Further, let $C \in L(W^1_A, X)$ be a symmetric linear operator in $X$ and $I$ be a non-empty open interval of $\mathbb{R}$.

Definition 4.5: We define a solution space $S_I$ to consist of all differentiable $u : I \to W^1_A$ with $\text{Ran}(u) \subset D(A)$, such that $u' : I \to X$ is differentiable and

$$(u'(t) + iBu'(t) + (A + C)u(t) = 0 \quad (4.1)$$

for every $t \in I$.^30

Note that (4.1) contains two types of derivatives. Every first derivative of $u$ is to be understood in the sense of derivatives of $W^1_A$-valued functions, whereas every further derivative is to be understood in the sense of derivatives of $X$-valued functions. Unless otherwise indicated, this convention is also adopted in the subsequent part of this section. On the other hand, since the imbedding $W^1_A \hookrightarrow X$ is continuous, differentiability in the sense of $W^1_A$-valued functions also implies differentiability in the sense of $X$-valued functions, including agreement of the corresponding derivatives. In particular, every $u \in S_I$ also satisfies the equation

$$u'(t) + iBu'(t) + (A + C)u(t) = 0 \quad (4.2)$$

for every $t \in I$, where here all derivatives are to be understood in the sense of derivatives of $X$-valued functions. Further, note that the assumptions on $C$, in general, do not imply that $A + C$ is self-adjoint.

Remark 4.6: According to Theorem 4.11 in Ref. 4, see also Theorem 5.4.11 in Ref. 5, for every $t_0 \in I$, $\xi \in D(A)$, and $\eta \in W^1_A$, there is a uniquely determined corresponding $u \in S_I$ such that $u(t_0) = \xi$ and $u'(t_0) = \eta$. The proof uses methods from the theory of semigroups of operators. Independently, the uniqueness of such $u$ follows more elementary from energy estimates in part (iii) of the subsequent Lemma 4.7.

Parts (i) and (ii) of the subsequent Lemma 4.7 give a “conserved current” and a “conserved energy,” respectively, that are associated with solutions of (4.1). Part (iii) gives associated energy estimates, that, in particular, imply the uniqueness of the initial value problem for (4.1) stated in (iv).

Lemma 4.7: Let $u \in S_I$ and $t_0 \in I$. Then the following holds.

(i) If $v \in S_I$, then $j_{u,v} : I \to \mathbb{C}$, defined by

$$j_{u,v}(t) := \langle u(t) | v'(t) \rangle - \langle u'(t) | v(t) \rangle + i \langle u(t) | Bv(t) \rangle$$

for every $t \in I$, is constant.

(ii) The function $E_u : I \to \mathbb{R}$, defined by

$$E_u(t) := \|u'(t)\|^2 + \langle u(t) | (A + C)u(t) \rangle$$

for every $t \in I$, is constant.

(iii) In addition, let $A + C$ be semibounded with lower bound $\gamma \in \mathbb{R}$. Then

$$\|u(t_2)\| \leq \begin{cases} \|u(t_1)\| + |E_u|^{1/2} (t_2 - t_1) & \text{if } \gamma < 0, \\ \|u(t_1)\| + E_u^{1/2} (t_2 - t_1) & \text{if } \gamma = 0, \\ (2E_u/\gamma)^{1/2} \left(1 - e^{-\gamma^{1/2}(t_2 - t_1)}\right) + \|u(t_1)\| e^{-\gamma^{1/2}(t_2 - t_1)} & \text{if } \gamma > 0, \end{cases}$$

for $t_1, t_2 \in I$ such that $t_1 \leq t_2$. 

(iv) In addition, let $A + C$ be semibounded. If $v \in S_I$ is such that

$$u(t_0) = v(t_0), \quad u'(t_0) = v'(t_0),$$

then $v = u$. 
The following example proves that it is possible that the energy assumes strictly negative values, but that the solutions of (4.1) are stable, i.e., that there are no exponentially growing solutions. This is different from the case of vanishing $B$, where there are unstable solutions of (4.1) if and only if the energy assumes strictly negative values.

**Example 4.8:** The example uses for the Hilbert space $X$ the space $\mathbb{C}^2$ equipped with the Euclidean scalar product, $\tilde{A} := A + C$ and $B$ are the linear operators on $\mathbb{C}^2$ whose representations with respect to the canonical basis are given by the matrices

\[
\begin{pmatrix}
1 & 0 \\
0 & -1
\end{pmatrix}
\quad \text{and} \quad
\begin{pmatrix}
3 & 1 \\
1 & 3
\end{pmatrix},
\]

(4.3)

respectively. An analysis shows that $\tilde{A}$ and $B$ are bounded linear and self-adjoint operators in $X$, $\tilde{A}$ is semibounded, $B$ is positive, and $\tilde{A} + (1/4)B^2$ is strictly positive. Further, $\tilde{A}$ and $B$ do not commute. Finally, the operator polynomial ($\mathbb{C} \rightarrow L(X,X), \lambda \mapsto \tilde{A} - \lambda B - \lambda^2$) has four distinct real eigenvalues. Therefore, in this case, there are no exponentially growing solutions of the corresponding equation (4.1). Figure 1 gives the graph of $p := (\mathbb{R} \rightarrow \mathbb{R}, \lambda \mapsto \det(\tilde{A} - \lambda B - \lambda^2)) = \lambda^4 + 6\lambda^3 + 8\lambda^2 - 1$ which suggests that there are precisely four distinct real roots. Indeed, we notice that

\[
p(-5) > 0, \quad p(-4) < 0, \quad p(-1) > 0, \quad p(0) < 0, \quad p(1) > 0
\]

and hence that $p$ has real roots in the intervals $(-5, -4)$, $(-4, -1)$, $(-1, 0)$, and $(0, 1)$. In addition, the value of the conserved energy $E_u$ corresponding to the solution $u$ of (4.1) with initial data $u(0) \equiv '0, 1)$ and $u'(0) \equiv '0, 0)$ is $<0$.

There are other possible definitions for the energy that is associated with solutions of (4.1). In cases of vanishing $B$, such are usually not of further use. In the case of a nonvanishing $B$, they can be useful as is the case for the RKG. In this case, the positivity of $E_{s, u}$ for sufficiently large masses of the field and

\[
s = \frac{ma}{2Mr_+}
\]

(4.4)

provides a basis for (1.3) and its improvement (4.5) below.
Corollary 4.9: Let \( s \in \mathbb{R} \) and \( u \in S_I \). Then, the function \( E_{s,u} : I \to \mathbb{R} \), defined by
\[
E_{s,u}(t) := \|u'(t) + isu(t)\|^2 + |u(t)|(A + C + s(B - s))u(t)
\]
for every \( t \in I \), is constant. If \( A + C + s(B - s) \) is additionally semibounded with lower bound \( \gamma \in \mathbb{R} \), then
\[
\|u(t_2)\| \leq \begin{cases} 
\|u(t_1)\| + |E_{s,u}|^{1/2}(t_2 - t_1) & \text{if } \gamma < 0, \\
\|u(t_1)\| + E_{s,u}^{1/2}(t_2 - t_1) & \text{if } \gamma = 0, \\
(2E_{s,u}/\gamma)^{1/2} \left(1 - e^{-\gamma/2(t_2 - t_1)}\right) + \|u(t_1)\|e^{-\gamma/2(t_2 - t_1)} & \text{if } \gamma > 0,
\end{cases}
\]
for \( t_1, t_2 \in I \) such that \( t_1 \leq t_2 \).

**Proof:** See Appendix B.

Theorem 4.10: If there is \( s \in \mathbb{R} \) such that \( A + C + s(B - s) \) is positive, then there are no exponentially growing solutions of (4.1).

**Proof:** The statement is a direct consequence of Corollary 4.9 (or Theorem 4.17 (ii) in Ref. 4, see also Theorem 5.4.17 (ii) in Ref. 5).

Assumption 4.11: In the following, we assume that \( X \) is given by (3.1), \( A := \tilde{A}_0 - C, B \) is given by (3.2), and \( C := -(\alpha + \epsilon) \) for some \( \epsilon > 0 \).

Theorem 4.10 leads to an improvement of the estimate (1.3).

**Theorem 4.12:** If
\[
\mu \geq \frac{|m|a}{2Mr_+} \sqrt{1 + \frac{2M}{r_+}}, \tag{4.5}
\]
then there are no exponentially growing solutions of (4.1).

**Proof:** Let \( s \in \mathbb{R} \). In the following, we estimate \( \tilde{A}_0 + sB - s^2 \). For this, let \( f \in D(A_0) \). Then
\[
(A_0 + sB - s^2)f = \frac{1}{\Sigma_b} \left( -\frac{\partial}{\partial r} \Delta \frac{\partial}{\partial r} - \frac{1}{\sin \theta} \frac{\partial}{\partial \theta} \sin \theta \frac{\partial}{\partial \theta} + \frac{m^2}{\sin^2 \theta} + V_s \right) f,
\]
where
\[
V_s := -\frac{m^2a^2}{\Delta} + \mu^2 \Sigma_b + s \frac{4mMr}{\Delta} - s^2 \Sigma_b
\]
\[
= -\frac{(2sMr - ma)^2}{\Delta} + (\mu^2 - s^2) \Sigma - 2s^2Mr.
\]
First, we note that
\[
\frac{m^2}{\sin^2 \theta} \geq m^2.
\]
In the following, we assume that \( s = ma/(2Mr_+) \). Then
\[
V_{s1} := -\frac{(2sMr - ma)^2}{\Delta} = - \left( \frac{ma}{r_+} \right)^2 + \left( \frac{ma}{r_+} \right)^2 \frac{2\sqrt{M^2 - a^2}}{r - r_+} \geq -m^2.
\]
Further, we define
\[
V_{s2} := (\mu^2 - s^2) \Sigma - 2s^2Mr = (\mu^2 - s^2)w^2 - 2s^2Mr + a^2(\mu^2 - s^2) \cos^2 \theta.
\]
If \( \mu \geq |s| \cdot \left( 1 + (2Mr_+) \right)^{1/2} \), then
\[
V_{s2} \geq s^2 \frac{2M}{r_+} r^2 - 2s^2Mr + a^2(\mu^2 - s^2) \cos^2 \theta \geq a^2(\mu^2 - s^2) \cos^2 \theta \geq 0.
\]
As a consequence,
\[ \frac{1}{\Sigma b} \left( \frac{m^2}{\sin^2 \theta} + V_s \right) \geq 0. \]

Further, we conclude that
\[ \langle f \otimes (P^m_1 \circ \cos) | (A_0 + sB - s^2)(f \otimes (P^m_1 \circ \cos)) \rangle \]
\[ \geq \int_{\Omega_1} \sin \theta \left( f \otimes (P^m_1 \circ \cos) \right)^* \left( -\frac{\partial}{\partial r} \Delta \frac{\partial}{\partial r} - \frac{1}{\sin \theta} \frac{\partial}{\partial \theta} \sin \theta \frac{\partial}{\partial \theta} \right) \]
\[ (f \otimes (P^m_1 \circ \cos)) dr \, d\theta \geq 0 \]
for every \( f \in C^2_0((r_+^2, \infty), \mathbb{C}) \) and \( l \in \{|m|, |m| + 1, \ldots\} \). Since \( D \) is a core for \( \bar{A}_0 \), this implies that
\[ \bar{A}_0 + sB - s^2 \geq 0. \]
Hence the statement follows from Theorem 4.10.

The following gives a connection of the operator \( \bar{A}_0 + sB - s^2, s \in \mathbb{R} \), and the Killing field \( \partial_t + s \partial \phi \). The corresponding proof is given in Appendix B. This connection sheds light on the previous proof of the positivity of \( \bar{A}_0 + sB - s^2 \) for \( s = ma/(2Mr + \mu) \) for sufficiently large \( \mu \). Differently to \( g_{tt} \), the term \( g(\partial_t + s \partial \phi, \partial_t + s \partial \phi) \) is positive in a neighborhood of the event horizon, but gradually turns negative away from the horizon. The latter is compensated by the mass term \( \mu^2 \rho \) for sufficiently large \( \mu \).

**Lemma 4.13:** Let \( s \in \mathbb{R} \) and \( \xi := \partial_t + s \partial \phi \). Then
\[ \left[ A_0 + msB - (ms)^2 \right] f \]
\[ = \frac{1}{g^{tt}} \left[ \frac{1}{\sqrt{-|g|}} \partial_r \sqrt{-|g|} g^{rr} \partial_r + \frac{1}{\sqrt{-|g|}} \partial_\theta \sqrt{-|g|} g^{\theta \theta} \partial_\theta \right] f + \frac{m^2 g(\xi, \xi) + \mu^2 \rho}{-g_{\phi \phi}} f \]
for every \( f \in D(A_0) \), where
\[ \rho := -[g_{tt} g_{\phi \phi} - (g_{\phi \phi})^2] = \Delta \sin^2 \theta. \]

**Proof:** See Appendix B.

Subsequently, we rewrite (4.2) into an equivalent time-dependent wave equation that is governed by a family of unitarily equivalent positive self-adjoint operators. The latter equation might turn out useful for further investigation since only self-adjoint operators are involved. On the other hand, a subsequent example proves that from the abstract properties of this family alone it cannot be concluded that the solutions of the equation are stable.

**Lemma 4.14:** Let \( B \) be additionally bounded and \( u \in S_f \). Then, \( v : I \to X \) defined by
\[ v(t) := \exp((it/2)B)u(t) \]
for every \( t \in I \) is twice differentiable in the sense of derivatives of \( X \)-valued functions and satisfies
\[ v''(t) + A(t)v(t) = 0 \quad (4.6) \]
for every \( t \in I \), where
\[ A(t) := \exp((it/2)B) \left( A + C + \frac{1}{4} B^2 \right) \exp(-i(t/2)B) \quad (4.7) \]
for every \( t \in \mathbb{R} \).

**Proof:** See Appendix B.
The previous can be used to prove the stability of the solutions of (4.1) in particular cases where the operators $A + C$ and $B$ commute. Note that in these cases, there is a further conserved “energy” associated to the solutions of (4.1).

**Theorem 4.15:** If, in addition, $A + C$ is self-adjoint and semibounded, $B$ is bounded, $A + C$ and $B$ commute, i.e.,

$$(A + C) \circ B \supset B \circ (A + C)$$

and

$$A + C + \frac{1}{4} B^2$$

is positive, then there are no exponentially growing solutions of (4.1).

**Proof:** The statement is a simple consequence of Lemma 4.14 and Lemma 4.7 (iii).

Coming back to the statement of Lemma 4.14, for every $t \in I$, the corresponding $A(t)$ is a densely defined, linear, and self-adjoint operator in $X$, see, e.g., Lemma 5.1, in the Appendix. In particular, if $A + C + (1/4) B^2$ is positive, $A(t)$ is positive, too. For instance, according to Lemma 3.7, this is true in the special case of the Klein-Gordon equation (2.1). Hence, in such case it might be expected that (4.6) for $u \in S_t$ implies that $|u|$ is not exponentially growing since this is the case if $A(t) = A$ for every $t \in I$, where $A$ is a densely defined, linear, positive self-adjoint operator in $X$. In that case, $u$ is given by

$$u(t) = \cos((t - t_0)A^{1/2})u(t_0) + \frac{\sin((t - t_0)A^{1/2})}{A^{1/2}} u'(t_0)$$

for all $t_0, t \in I$, where $\cos((t - t_0)A^{1/2})$ and $\sin((t - t_0)A^{1/2})$ denote the bounded linear operators that are associated with the functional calculus for $A^{1/2}$ to the restriction of $\cos((t - t_0)\text{id}_X)$ and the restriction of the continuous extension of $\sin((t - t_0)\text{id}_X)\text{id}_X$ to $[0, \infty)$, respectively, to the spectrum of $A^{1/2}$. Note that the solutions (4.8) are in particular bounded if $A$ is strictly positive. Unfortunately, this expectation is in general not true. A counterexample can be found already on the level of finite-dimensional Hilbert spaces.

**Example 4.16:** The example uses for the Hilbert space $X$ the space $\mathbb{C}^2$ equipped with the Euclidean scalar product, $\tilde{A} := A + C$ and $B$ are the linear operators on $\mathbb{C}^2$ whose representations with respect to the canonical basis are given by the matrices

$$\begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} \quad \text{and} \quad \begin{pmatrix} 23/10 & 1 \\ 1 & 23/10 \end{pmatrix},$$

(4.9)

respectively. An analysis shows that $\tilde{A}$ and $B$ are bounded linear and self-adjoint operators in $X$, $\tilde{A}$ is semibounded, $B$ is positive, and $\tilde{A} + (1/4) B^2$ is even strictly positive. Further, $\tilde{A}$ and $B$ do not commute. Finally, the operator polynomial $(C \rightarrow L(X, X), \lambda \mapsto \tilde{A} - \lambda B - \lambda^2)$ has an eigenvalue with real part $< 0$. Therefore, in this case, there is an exponentially growing solution of the corresponding equation (4.2) and hence also of (4.6). Note that in this case, the corresponding family of operators (4.7) consists of strictly positive bounded self-adjoint linear operators whose spectra are bounded from below by a common strictly positive real number. Figure 2 gives the graph of $p := (\mathbb{R} \rightarrow \mathbb{R}, \lambda \mapsto \det(\tilde{A} - \lambda B - \lambda^2)) = \lambda^4 + 4.6\lambda^3 + 4.29\lambda^2 - 1$ which suggests that there are precisely two distinct simple roots. Indeed, this is true. The proof proceeds by a discussion of the graph of $p$ using the facts that

$$p(-4) > 0, \quad p(-3) < 0, \quad p(0) < 0, \quad p(1) > 0,$$

that the zeros of $p'$ are given by

$$(-69 - \sqrt{1329})/40, \quad (-69 + \sqrt{1329})/40, \quad 0$$

and that

$$p((-69 + \sqrt{1329})/40) < 0.$$
Thus, \((C \rightarrow L(X, X), \lambda \mapsto \det(\tilde{A} - \lambda B - \lambda^2))\) has two distinct simple real roots and a pair of simple complex conjugate roots.

V. DISCUSSION

The mathematical investigation of the stability of Kerr space-time has started but is still in the phase of the study of relevant model equations in a Kerr background. The study of the solutions of the Klein-Gordon equation is expected to give important insight into the problem.

In the case of the wave equation, i.e., for the case of vanishing mass \(\mu\) of the scalar field, results point to the stability of the solutions. On the other hand, inspection of the reduced Klein-Gordon equation, (2.1), reveals that the case of \(\mu > 0\) originates from the case \(\mu = 0\) by the addition of a positive bounded potential term

\[
\mu^2 \frac{\Sigma}{\Sigma_b}
\]

to the equation. If there were no first order time derivative present in the equation, from this alone, it would be easy to prove that the stability of the solutions of the wave equation implies the stability of the solutions of the Klein-Gordon equation for non-vanishing mass.

In presence of such a derivative, such stabilizing influence of a bounded positive potential term is not clear, but it is worth noting that the energy estimates in Lemma 4.7 indicate a stabilizing influence of such a term.

Interestingly, according to an unknown referee, there is another range of masses for which stability results can be shown. The method of Dafermos and Rodnianski from Ref. 13 can be adapted to show the stability of the solutions for masses \(\mu\) below some bound that depends on the azimuthal separation parameter \(m\).

The numerical results that indicate instability in the case \(\mu \neq 0\) make quite special assumptions on the values of the rotational parameter of the black hole that do not make them look very trustworthy. They could very well be numerical artifacts. Moreover, the numerical investigation by Konoplya and Zhidenko\(^{25}\) does not find any unstable modes and contradicts all these investigations. Also the analytical results in this area are not accompanied by error estimates and, therefore, ultimately inconclusive. Still, apart from Ref. 42, all these results are consistent with the estimate on \(\mu\) in Ref. 3 and the improved estimate of this paper, above which the solutions of the reduced, by separation in the azimuth angle in Boyer-Lindquist coordinates, Klein-Gordon equation are stable.
It seems that the proof of the stability of the solutions of the wave equation in a Kerr background will soon be established. The question of the stability of the massive scalar field in a Kerr background is still an open problem, with only few rigorous results available, and displays surprising mathematical subtlety. In particular, in this case standard tools of theoretical physical investigation, including numerical investigations, seem too imprecise for analysis. Hence a rigorous mathematical investigation, like the one performed in this paper, seems to be enforced.

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APPENDIX A: PROOFS FROM SECTION II

In the following, we give the proofs of the Lemmas 2.1 and 2.2 from Sec. II.

Proof of Lemma 2.1.

Proof: For this, let $s \in \mathbb{R}$. Then

$$g(\partial_t + s \partial \varphi, \partial_t + s \partial \varphi) = g_{tt} + 2s g_{t \varphi} + s^2 g_{\varphi \varphi}$$

$$= 1 - \frac{2Mr}{\Sigma} + 4s \frac{Mr \sin^2 \theta}{\Sigma} - \frac{s^2 \Delta \Sigma \rho}{\Sigma} \sin^2 \theta$$

$$= \frac{\Delta}{\Sigma} + \frac{\sin^2 \theta}{\Sigma} \left[-a^2 + 4s Mr - s^2 (r^2 + a^2)^2 + s^2 \Delta \sin^2 \theta \right]$$

$$= \frac{\Delta}{\Sigma} + \frac{\sin^2 \theta}{\Sigma} \left[-(a - 2s Mr)^2 + 4s^2 Mr^2 - s^2 (r^2 + a^2)^2 + a^2 s^2 \Delta \sin^2 \theta \right]$$

$$= \frac{\Delta}{\Sigma} - \frac{\sin^2 \theta}{\Sigma} \left[(a - 2s Mr)^2 + s^2 \Delta (\Delta + 4Mr - a^2 \sin^2 \theta) \right].$$

Hence, $g(\partial_t + s \partial \varphi, \partial_t + s \partial \varphi)$ has a positive extension to the boundary of $\Omega$, if and only if

$$s = \frac{a}{2Mr^+}.$$

In this case,

$$\left(a - 2s Mr\right)^2 + s^2 \Delta (\Delta + 4Mr - a^2 \sin^2 \theta)$$

$$= \frac{a^2}{r^+} (r - r_+)^2 + \frac{a^2}{4Mr^2} \Delta (\Delta + 4Mr - a^2 \sin^2 \theta)$$

$$= \frac{a^2}{4Mr^2} \left[4M^2 (r - r_+)^2 + \Delta (\Delta + 4Mr - a^2 \sin^2 \theta) \right]$$

$$= \frac{a^2 \Delta}{4Mr^2} \left[4M^2 \frac{r - r_+}{r - r_-} + \Delta + 4Mr - a^2 \sin^2 \theta \right].$$
and hence
\[ g(\partial_t + s \partial_{\varphi}, \partial_t + s \partial_{\varphi}) \]
\[ = \frac{\Delta}{4M^2 r_+^4 \Sigma} \left[ 4M^2 r_f^2 - a^2 \sin^2 \theta \left( 4M^2 \frac{r-r_+}{r-r_-} + \Delta + 4Mr - a^2 \sin^2 \theta \right) \right] \]
\[ = \frac{\Delta}{4M^2 r_+^4 \Sigma} \left[ (2Mr_+ - a^2 \sin^2 \theta)^2 - a^2 (r-r_+) \sin^2 \theta \left( \frac{4M^2}{r-r_-} + r - r_+ + 4M \right) \right] \]
\[ = \frac{\Delta}{4M^2 r_+^4 \Sigma} \left[ (2Mr_+ - a^2 \sin^2 \theta)^2 - a^2 \Delta \sin^2 \theta \left( 1 + \frac{2M}{r-r_-} \right)^2 \right]. \]

\[ \square \]

**Proof of Lemma 2.2.**

*Proof:* For this, let \((r, \theta) \in \Omega_{c1} . \) Then
\[ \Delta(r, \theta) < a^2 \sin^2 \theta \]
and
\[ (2Mr_+ - a^2 \sin^2 \theta)^2 - a^2 \Delta(r, \theta) \sin^2 \theta \left( 1 + \frac{2M}{r-r_-} \right)^2 \]
\[ = a^4 \sin^4 \theta - \left[ 4Mr_+ + \Delta(r, \theta) \left( 1 + \frac{2M}{r-r_-} \right)^2 \right] a^2 \sin^2 \theta + 4M^2 r_+^2 \]
\[ > (\Delta(r, \theta))^2 - \left[ 4Mr_+ + \Delta(r, \theta) \left( 1 + \frac{2M}{r-r_-} \right)^2 \right] a^2 + 4M^2 r_+^2 \]
\[ = (\Delta(r, \theta))^2 - a^2 \Delta(r, \theta) \left( 1 + \frac{2M}{r-r_-} \right)^2 + 4Mr_+ (Mr_+ - a^2) \]
\[ = \left[ \Delta(r, \theta) - \frac{a^2}{2} \left( 1 + \frac{2M}{r-r_-} \right)^2 \right] - \frac{a^4}{4} \left( 1 + \frac{2M}{r-r_-} \right)^4 + 4Mr_+ (Mr_+ - a^2) \]
\[ \geq \left[ \Delta(r, \theta) - \frac{a^2}{2} \left( 1 + \frac{2M}{r-r_-} \right)^2 \right] + 4 \left[ - \frac{a^4 r_+^4}{(r_+ - r_-)} + Mr_+ (Mr_+ - a^2) \right]. \]

Hence it follows that \((r, \theta) \in \Omega_{c2} \) if
\[ \frac{a^4 r_+^4}{(r_+ - r_-)^4} + a^2 Mr_+ - M^2 r_+^2 \]
\[ = \frac{r_+^4}{(r_+ - r_-)^4} \left[ a^4 + \frac{M(r_+ - r_-)^4}{r_+^3} a^2 - \frac{M^2 (r_+ - r_-)^4}{r_+^2} \right] \leq 0. \]

The latter is the case if and only if
\[ a^2 \leq \frac{2Mr_+}{1 + \sqrt{1 + \frac{4r_+^4}{(r_+ - r_-)^4}}}. \]

Further,
\[ \frac{2Mr_+}{1 + \sqrt{1 + \frac{4r_+^4}{(r_+ - r_-)^4}}} \geq \frac{Mr_+}{1 + \frac{r_+^2}{(r_+ - r_-) \sqrt{a^2}}} \geq \frac{M^2}{1 + \frac{M^2}{M^2 - a^2}} = \frac{M^2 (M^2 - a^2)}{2M^2 - a^2} \]
\[ \geq \frac{1}{2} (M^2 - a^2). \]
Hence if
\[ a^2 \leq \frac{1}{2} (M^2 - a^2), \]
or, equivalently, if condition (2.2) is satisfied, it follows that \((r, \theta) \in \Omega_2\). 
\(\square\)

**APPENDIX B: PROOFS FROM SECTIONS III AND IV**

In the following, we give the omitted proofs from Secs. III and IV.

**Proof of Lemma 4.7.**

Proof: “(i)”: For this, let \( t \in I \) and \( h \in \mathbb{R} \) such that \( t + h \in I \). Then
\[
\frac{j_{u,v}(t + h) - j_{u,v}(t)}{h} = \frac{h^{-1} \left[ (u(t + h)|v'(t + h)) - (u'(t + h)|v(t + h)) + i \langle u(t + h)|Bu(t + h) \rangle \right.}{h} \\
= \frac{h^{-1} \left[ (u(t + h)|v'(t + h)) - (u'(t)|v(t)) - i \langle u(t)|Bu(t) \rangle \right.}{h}
\]
Hence it follows that \( j_{u,v} \) is differentiable in \( t \) with derivative
\[
j'_{u,v}(t) = (u'(t)|v'(t)) - (u'(t)|v(t)) + i \langle u'(t)|Bu(t) \rangle \]
\[
\quad = (u'(t)|v'(t)) + iBv(t) - (u'(t) + iBu(t))v(t) \\
\quad = -(u(t)|Av(t)) + (A+C)u(t)v(t) = 0.
\]

From the latter, we conclude that the derivative of \( j_{u,v} \) vanishes and hence that \( j_{u,v} \) is a constant function.

“(ii)”: For this, again, let \( t \in I \) and \( h \in \mathbb{R} \) such that \( t + h \in I \). Further, let \( \tilde{A} := A + C \). Then
\[
\frac{E_u(t + h) - E_u(t)}{h} = \frac{h^{-1} \left[ (u'(t + h)|u'(t + h)) + (u(t + h)|\tilde{A}u(t + h)) - (u'(t)|u'(t)) - (u(t)|\tilde{A}u(t)) \right.}{h}
\]
Hence it follows that \( E_u \) is differentiable in \( t \) with derivative
\[
(u'(t)|u'(t)) + (u'(t)|u'(t)) + (A^{1/2}u(t + h)) + (u(t + h) - u(t)Cu(t + h)) \\
\quad = -(iBu(t) + (A + C)u(t)|u'(t)) - (u'(t)|Bu'(t) + (A + C)u(t)) \\
\quad = -(iBu(t)|u'(t)) + (u'(t)|Bu'(t)) = 0.
\]

From the latter, we conclude that the derivative of \( E_u \) vanishes and hence that \( E_u \) is a constant function.
“(iii)”: Since $A + C$ is semibounded with lower bound $\gamma \in \mathbb{R}$,
\[
\langle \xi | (A + C)\xi \rangle \geq \gamma \|\xi\|^2
\]
for every $\xi \in D(A)$. Hence it follows by (ii) that
\[
\|u'(t)\|^2 + \gamma \|u(t)\|^2 = E_u - (u(t)(A + C)u(t)) - \gamma \|u(t)\|^2 \leq E_u
\]
for every $t \in \mathbb{R}$. If $\gamma = 0$, the latter implies that
\[
\|u'(t)\| \leq E_u^{1/2}
\]
for every $t \in I$. Hence it follows by weak integration in $X$, e.g., see Theorem 3.2.5 in Ref. 5, that
\[
\|u(t_2) - u(t_1)\| = \left\| \int_{(t_1,t_2)} u'(t) dt \right\| \leq \int_{(t_1,t_2)} \|u'(t)\| dt \leq E_u^{1/2}(t_2 - t_1),
\]
where $t_1, t_2 \in I$ are such that $t_1 < t_2$, and hence that
\[
\|u(t_2)\| \leq \|u(t_1)\| + E_u^{1/2}(t_2 - t_1).
\]
For the weak integration, note that the inclusion of $W_2^1$ into $X$ is continuous. If $\gamma > 0$, it follows from (B1) along with the parallelogram identity for elements of $X$ that
\[
\|e^{-\gamma^{1/2}}(e^{\gamma^{1/2} \text{id}_{\mathbb{R}}^n} u)'(t)\|^2 = \|u'(t) + \gamma^{1/2}u(t)\|^2 \leq 2(\|u'(t)\|^2 + \|\gamma^{1/2}u(t)\|^2) \leq 2E_u
\]
and hence that
\[
\|(e^{\gamma^{1/2} \text{id}_{\mathbb{R}}^n} u)'(t)\| \leq (2E_u)^{1/2} e^{\gamma^{1/2} t}
\]
for $t \in I$. Hence it follows by weak integration in $X$ that
\[
\|e^{\gamma^{1/2} t}u(t_2) - e^{\gamma^{1/2} t_1}u(t_1)\| = \left\| \int_{(t_1,t_2)} (e^{\gamma^{1/2} \text{id}_{\mathbb{R}}^n} u)'(t) dt \right\|
\leq \int_{(t_1,t_2)} \|(e^{\gamma^{1/2} \text{id}_{\mathbb{R}}^n} u)'(t)\| dt \leq (2E_u/\gamma)^{1/2} \left(e^{\gamma^{1/2} t_2} - e^{\gamma^{1/2} t_1}\right)
\]
for all $t_1, t_2 \in I$ such that $t_1 < t_2$. The latter implies that
\[
\|e^{\gamma^{1/2} t}u(t_2)\| \leq \|e^{\gamma^{1/2} t_1}u(t_1)\| + (2E_u/\gamma)^{1/2} \left(e^{\gamma^{1/2} t_1} - e^{\gamma^{1/2} t_2}\right).
\]
Hence
\[
\|u(t_2)\| \leq (2E_u/\gamma)^{1/2} \left(1 - e^{-\gamma^{1/2}(t_2-t_1)}\right) + e^{-\gamma^{1/2}(t_2-t_1)}\|u(t_1)\|.
\]
If $\gamma < 0$, it follows from (B1) that
\[
\|u'(t)\|^2 \leq E_u - \gamma \|u(t)\|^2 \leq |E_u| + a \|u(t)\|^2,
\]
for every $t \in I$, where $a := -\gamma > 0$. The latter implies that
\[
\|u'(t)\| \leq |E_u|^{1/2} + a^{1/2} \|u(t)\|
\]
for every $t \in I$. Hence it follows by weak integration in $X$ that
\[
\|u(t_2) - u(t_1)\| = \left\| \int_{(t_1,t_2)} u'(t) dt \right\| \leq \int_{(t_1,t_2)} \|u'(t)\| dt
\leq |E_u|^{1/2}(t_2 - t_1) + a^{1/2} \int_{(t_1,t_2)} \|u(t)\| dt,
\]
where $t_1, t_2 \in I$ are such that $t_1 < t_2$, and
\[
\|u(t_2)\| \leq \|u(t_1)\| + |E_u|^{1/2}(t_2 - t_1) + a^{1/2} \int_{(t_1,t_2)} \|u(t)\| dt.
\]
By help of the generalized Gronwall inequality from Lemma 3.1 in Ref. 20, from the latter we conclude that

$$\|u(t_2)\| \leq [\|u(t_1)\| + |E_u|^{1/2}(t_2 - t_1)]e^{\gamma(t_2 - t_1)}$$

for $t_1 \in I$ and $t_2 \in I$ such that $t_1 < t_2$.

"(iv)" For this, we define $\nu := \nu - u$. Then $\nu$ is an element of $S_I$ such that $\nu(t_0) = \nu'(t_0) = 0$. This implies that

$$E_u(t) := \|\nu(t)\|^2 + (\nu(t)(A + C)u(t))$$

for every $t \in I$ is constant of value 0. Hence we conclude from (iii) that $\nu(t) = 0$ for all $t \in I$ and therefore that $\nu = u$.

\[\square\]

**Proof of Corollary 4.9.**

**Proof:** We define $v : I \to W^1_4$ by

$$v(t) := e^{is}u(t)$$

for every $t \in I$. Then $v$ is differentiable with $\text{Ran} \, v \subset D(A)$ and also $v' : I \to X$ is differentiable such that

$$(v')'(t) = e^{is}[u'(t) + isu(t)], \quad (v'')(t) = e^{is}[(u'')'(t) + 2isu'(t) - s^2u(t)]$$

for every $t \in I$. Further,

$$(v')'(t) + i(B - 2s)v'(t) + (A + C + sB - s^2)v(t)$$

$$\quad = e^{is}[u'(t) + 2isu(t) - s^2u(t) + i(B - 2s)(u'(t) + isu(t))$$

$$\quad \quad + (A + C + sB - s^2)u(t)]$$

$$\quad = e^{is}[u'(t) + 2isu(t) - s^2u(t) + IBu'(t) - 2isu(t) - Bu(t) + 2s^2u(t)$$

$$\quad \quad + (A + C + sB - s^2)u(t)]$$

$$\quad = e^{is}[u'(t) + IBu'(t) + (A + C)u(t)]$$

for every $t \in I$. Note that $(X, A, B - 2s, C + sB - s^2)$ satisfy Assumptions 4.1 and 4.4. Hence it follows by Lemma 4.7 that the function $E_v : I \to \mathbb{R}$, defined by

$$E_v(t) := \|v(t)\|^2 + (\nu(t)(A + C + sB - s^2)v(t))$$

$$\quad = \|u'(t) + isu(t)\|^2 + (u(t)(A + C + sB - s^2)u(t))$$

for every $t \in I$, is constant. If, in addition, $A + C + s(B - s)$ is semibounded with lower bound $\gamma \in \mathbb{R}$, then

$$\|u(t_2)\| \leq \begin{cases} \|u(t_1)\| + |E_v|^{1/2}(t_2 - t_1)[e^{\gamma/2}(t_2 - t_1)] & \text{if } \gamma < 0, \\ \|u(t_1)\| + E_v^{1/2}(t_2 - t_1) & \text{if } \gamma = 0, \\ (2E_v/\gamma)^{1/2}(1 - e^{-\gamma/2}(t_2 - t_1)) + \|v(t_1)e^{-\gamma/2}(t_2 - t_1) & \text{if } \gamma > 0, \end{cases}$$

for $t_1, t_2 \in I$ such that $t_1 \leq t_2$.

\[\square\]

**Proof of Lemma 4.13.**

**Proof:** First, we notice that the only non-vanishing components of $(g_{ab})_{(a, b) \in \{r, \theta, \phi\}}$ are given by

$$g^{rr} = \frac{\Sigma_b}{\Sigma}, \quad g^{\theta\theta} = g^{tt} = \frac{2Mar}{\Delta \Sigma}, \quad g^{rr} = \frac{\Delta}{\Sigma}, \quad g^{\theta\theta} = \frac{1}{\Sigma},$$

$$g^{\phi\phi} = \frac{1}{\Delta \sin^2 \theta} \left(1 - \frac{2Mr}{\Sigma}\right).$$
Further, we notice that
\[ g^{\nu} = - \frac{g_{\varphi \varphi}}{\rho}, \quad g^{\nu} = \frac{g_{\theta \theta}}{\rho}, \quad g^{\nu} = - \frac{g_{\nu}}{\rho}, \]
where
\[ \rho := - [g_{\nu}g_{\varphi \varphi} - (g_{\varphi})^2] = \Delta \sin^2 \theta. \]

Hence
\[ \frac{1}{g^{\nu}} \partial_\nu \partial_\nu = \partial_\nu^2 + 2 \frac{g^{\nu}}{g^{\mu}} \partial_\nu \partial_\mu + \frac{g^{\nu}}{g^{\varphi \varphi}} \partial_\varphi^2 \]
\[ + \frac{1}{g^{\mu}} \left[ \frac{1}{\sqrt{-|g|}} \partial_\mu \sqrt{-|g|} g^{\mu \varphi} \partial_\varphi + \frac{1}{\sqrt{-|g|}} \partial_\varphi \sqrt{-|g|} g^{\varphi \mu} \partial_\mu \right] \]
\[ = \partial_\nu^2 + 2 \frac{g_\nu}{-g_{\varphi \varphi}} \partial_\nu \partial_\varphi - \frac{g_\nu}{-g_{\varphi \varphi}} \partial_\varphi^2 \]
\[ + \frac{1}{g^{\mu}} \left[ \frac{1}{\sqrt{-|g|}} \partial_\mu \sqrt{-|g|} g^{\mu \varphi} \partial_\varphi + \frac{1}{\sqrt{-|g|}} \partial_\varphi \sqrt{-|g|} g^{\varphi \mu} \partial_\mu \right]. \]

As a consequence,
\[ A_0 f = \frac{1}{g^{\mu}} \left[ \frac{1}{\sqrt{-|g|}} \partial_\mu \sqrt{-|g|} g^{\mu \varphi} \partial_\varphi + \frac{1}{\sqrt{-|g|}} \partial_\varphi \sqrt{-|g|} g^{\varphi \mu} \partial_\mu \right] f + \frac{m^2 g_{\mu} + \mu^2 \rho}{-g_{\varphi \varphi}} f \]
for every \( f \in D(A_0) \). Finally, it follows that
\[ [A_0 + m s B - (m s)^2] f = A_0 f + m s 2m \frac{g^{\nu}}{g^{\mu}} f - (m s)^2 f \]
\[ = \frac{1}{g^{\mu}} \left[ \frac{1}{\sqrt{-|g|}} \partial_\mu \sqrt{-|g|} g^{\mu \varphi} \partial_\varphi + \frac{1}{\sqrt{-|g|}} \partial_\varphi \sqrt{-|g|} g^{\varphi \mu} \partial_\mu \right] f \]
\[ + \frac{m^2}{-g_{\varphi \varphi}} (g_{\mu} + 2 s g_{\varphi \varphi} + s^2 g_{\varphi \varphi}) f + \frac{\mu^2 \rho}{-g_{\varphi \varphi}} \]
\[ = \frac{1}{g^{\mu}} \left[ \frac{1}{\sqrt{-|g|}} \partial_\mu \sqrt{-|g|} g^{\mu \varphi} \partial_\varphi + \frac{1}{\sqrt{-|g|}} \partial_\varphi \sqrt{-|g|} g^{\varphi \mu} \partial_\mu \right] f + \frac{m^2 g(\xi, \xi) + \mu^2 \rho}{-g_{\varphi \varphi}} f \]
for every \( f \in D(A_0) \). \( \square \)

**Proof of Lemma 4.14.**

**Proof:** First, if \( D \in L(X, X) \) and \( f : I \to X \) is differentiable in \( t \in I \) and \( h \in \mathbb{R}^+ \) such that \( t + h \in I \), it follows that
\[ \frac{1}{h} \left[ \exp((t + h)D) f(t + h) - \exp(tD) f(t) \right] = \exp(tD) \frac{1}{h} \left[ \exp(hD) f(t + h) - f(t) \right] \]
\[ = \exp(tD) \left[ \exp(hD) \frac{1}{h} \left[ f(t + h) - f(t) \right] + \frac{1}{h} \left( \exp(hD) f(t) - f(t) \right) \right] \]
\[ = \exp(tD) \left[ \exp(hD) \left( \frac{1}{h} \left[ f(t + h) - f(t) \right] - f'(t) \right) + \exp(hD) f'(t) \right] \]
\[ + \frac{1}{h} \left( \exp(hD) f(t) - f(t) \right) \]
and hence that \( g := (I \to X, s \to \exp(sD)f(s)) \) is differentiable in \( t \) with derivative
\[ \exp(tD)[f'(t) + Df(t)]. \]
In particular, this implies, if \( f \) is twice differentiable in \( t \in I \), that \( g \) is twice differentiable in \( t \) with second derivative

\[
\exp(tD)[f''(t) + 2Df'(t) + D^2f(t)]
\]

Applying the previous auxiliary result to \( D = (i/2)B \) proves that \( v \) is twice differentiable. Further, from the definition of \( v \), it follows that

\[
u(t) = \exp(-(it/2)B)v(t)
\]

for every \( t \in I \). Application of the auxiliary results above to \( D = -(i/2)B \) leads to

\[
u'(t) = \exp(-(it/2)B)\left(v'(t) - \frac{i}{2}Bv(t)\right),
\]

\[
u''(t) = \exp(-(it/2)B)\left(v''(t) - iBv'(t) - \frac{1}{4}B^2v(t)\right).
\]

Hence it follows from (4.2) that

\[
0 = u''(t) + iBu'(t) + A*u(t)
\]

\[
= \exp(-(it/2)B)\left(v''(t) - iBv'(t) - \frac{1}{4}B^2v(t) + iBv'(t) - iB\frac{1}{2}Bv(t)
\]

\[+ \exp((it/2)B)\tilde{A}\exp(-(it/2)B)v(t)\right)
\]

\[
= \exp(-(it/2)B)\left(v''(t) + \frac{1}{4}B^2v(t) + \exp((it/2)B)\tilde{A}\exp(-(it/2)B)v(t)\right)
\]

\[
= \exp(-(it/2)B)\left[v'(t) + \exp((it/2)B)(\tilde{A} + \frac{1}{4}B^2)\exp(-(it/2)B)v(t)\right],
\]

where \( \tilde{A} := A + C \).

In the following, we give some abstract lemmas that are applied in the text. For the convenience of the reader, corresponding proofs are added.

**Lemma 5.1** Let \((X, \langle \cdot, \cdot \rangle)\) be a Hilbert space over \( \mathbb{K} \in \{ \mathbb{R}, \mathbb{C} \} \), \( A \) a densely defined, linear, and self-adjoint operator in \( X \) and \( U \in L(X, X) \) be unitary. Then, \( A_U := U\circ A\circ U^{-1} \) is a densely defined, linear, and self-adjoint operator in \( X \). Further, if \( D \subseteq D(A) \) is a core for \( A \), then \( U(D) \) is a core for \( U\circ A\circ U^{-1} \). Also, if \( A \) is positive, then \( U\circ A\circ U^{-1} \) is positive, too.

**Proof:** First, we note that \( D(U\circ A\circ U^{-1}) = U(D(A)) \). Since \( D(A) \) is dense in \( X \), for \( \xi \in X \), there is a sequence of \( \xi_1, \xi_2, \ldots \) of elements of \( D(A) \) such that

\[
\lim_{\nu \to \infty} \xi_\nu = U^{-1}\xi.
\]

Hence also

\[
\lim_{\nu \to \infty} U\xi_\nu = \xi.
\]

As a consequence, \( U\circ A\circ U^{-1} \) is densely defined. Also, as composition of linear maps, \( U\circ A\circ U^{-1} \) is linear. In addition, for \( \xi, \eta \in D(A) \), it follows that

\[
\langle U\xi|U \circ A \circ U^{-1}U\eta \rangle = \langle \xi|A\eta \rangle = \langle A\xi|\eta \rangle = \langle U \circ A \circ U^{-1}U\xi|U\eta \rangle
\]

and hence that \( U\circ A\circ U^{-1} \) is symmetric. Further, if \( \xi \in D((U\circ A\circ U^{-1})^*) \), then

\[
\langle (U \circ A \circ U^{-1})^*\xi|U\eta \rangle = \langle \xi|(U \circ A \circ U^{-1)U} \eta \rangle = \langle U^{-1}\xi|A\eta \rangle
\]

for every \( \eta \in D(A) \). Hence \( \xi \in U(D(A)) \), and

\[
\langle U^{-1}\xi|A\eta \rangle = \langle AU^{-1}\xi|\eta \rangle = \langle UAU^{-1}\xi|U\eta \rangle
\]
for every $\eta \in D(A)$. Since $U(D(A))$ is dense in $X$, this implies that $(U \circ A \circ U^{-1})^* \xi = UAU^{-1} \xi$. As a consequence,

$$U A U^{-1} \supset (U \circ A \circ U^{-1})^*.$$

Hence it follows that $U \circ A \circ U^{-1}$ is self-adjoint. Further, let $D \subseteq D(A)$ be a core for $A$. As a consequence, for every $\xi \in D(A)$ there is a sequence $\xi_1, \xi_2, \ldots$ in $D$ such that

$$\lim_{v \to \infty} \xi_v = \xi, \quad \lim_{v \to \infty} A\xi_v = A\xi.$$

Hence $U\xi_1, U\xi_2, \ldots$ is a sequence in $U(D)$ such that

$$\lim_{v \to \infty} U\xi_v = U\xi, \quad \lim_{v \to \infty} UAU^{-1}U\xi_v = UAU^{-1}U\xi.$$

Therefore, $U(D)$ is a core for $UAU^{-1}$. Finally, if $A$ is positive, it follows for $\xi \in D(A)$ that

$$(U\xi | U \circ A \circ U^{-1} U\xi) = (U\xi | U A\xi) = (\xi | A\xi) \geq 0$$

and hence also the positivity of $UAU^{-1}$. \hfill \Box

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18 For the first, see, for instance, Ref. 5. For the second, see, for instance, chapter 2 of Ref. 29.
22 If not otherwise indicated, the symbols $t, r, \theta, \varphi$ denote coordinate projections whose domains will be obvious from the context. In addition, we assume the composition of maps, which includes addition, multiplication, and so forth, always to be maximally defined. For instance, the sum of two complex-valued maps is defined on the intersection of their domains. Finally, we use Planck units where the reduced Planck constant $\hbar$, the speed of light in vacuum $c$, and the gravitational constant $\gamma$, all have the numerical value 1.


30 Note that the differentiability of $u$ implies that $Ran u' \subset W^1_0$.


37 That is, if $A$ is a densely-defined, linear, symmetric, and semibounded operator in some Hilbert space $X$ and $\gamma \in \mathbb{R}$, then the Friedrichs extension of $A + \gamma$, $(A + \gamma)_F$, and the sum of the Friedrichs extension of $A$, $A_F$, and $\gamma$ coincide, $(A + \gamma)_F = A_F + \gamma$.


39 $W^1_0$ may be regarded as a generalized Sobolev space.

